



# Special and General Relativity

Boud Roukema

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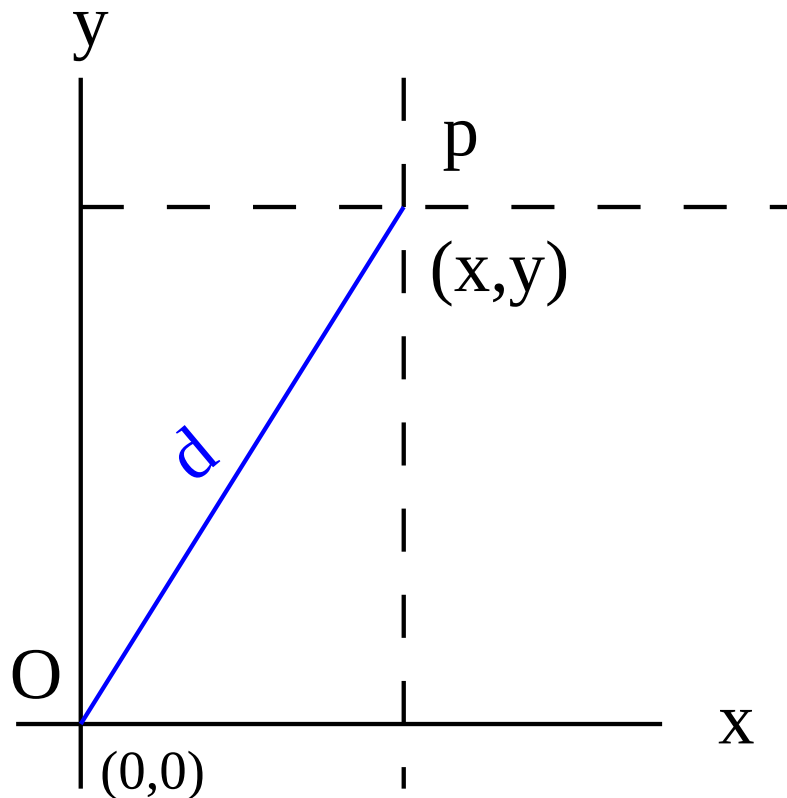
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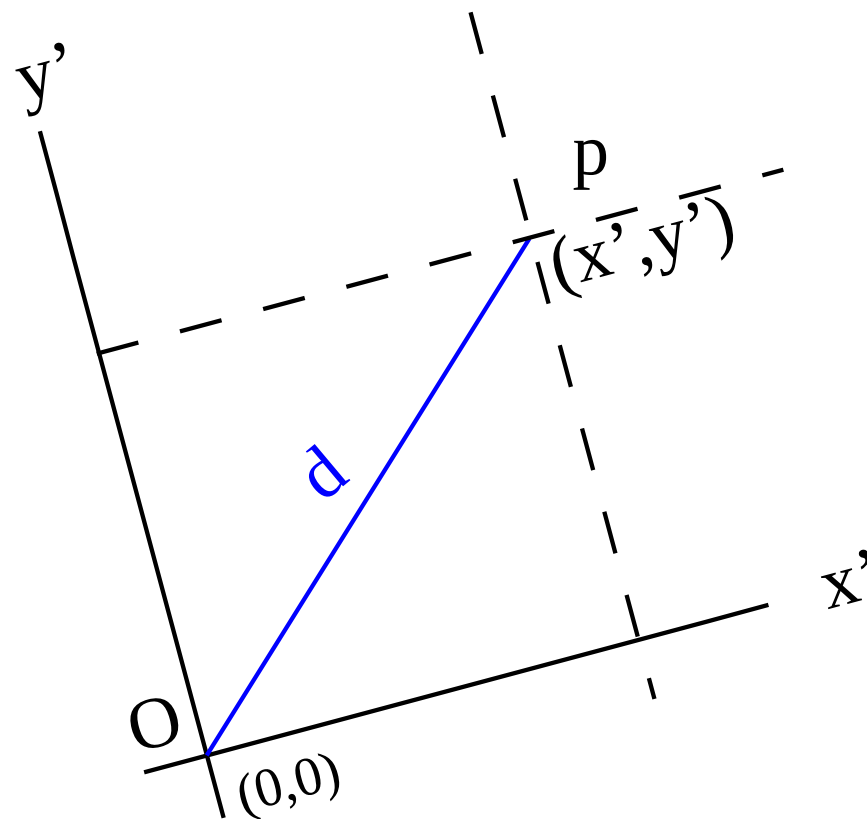
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- GR: spacetime = a solution of the w:Einstein field equations

# SR: Minkowski spacetime



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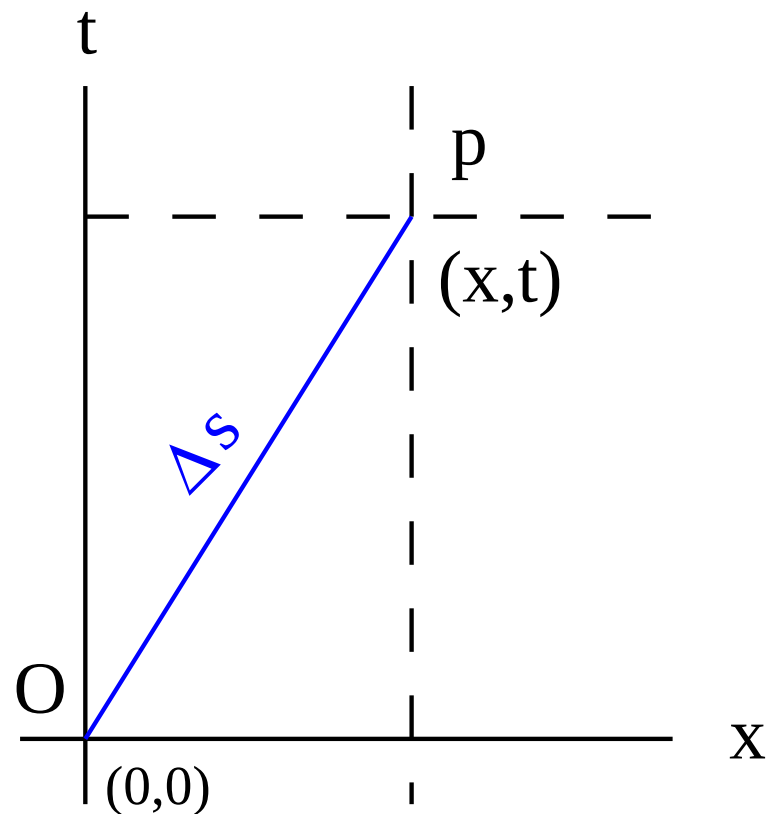
# SR: Minkowski spacetime



$$\begin{pmatrix} p_{x'} \\ p_{y'} \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} p_x \\ p_y \end{pmatrix}$$

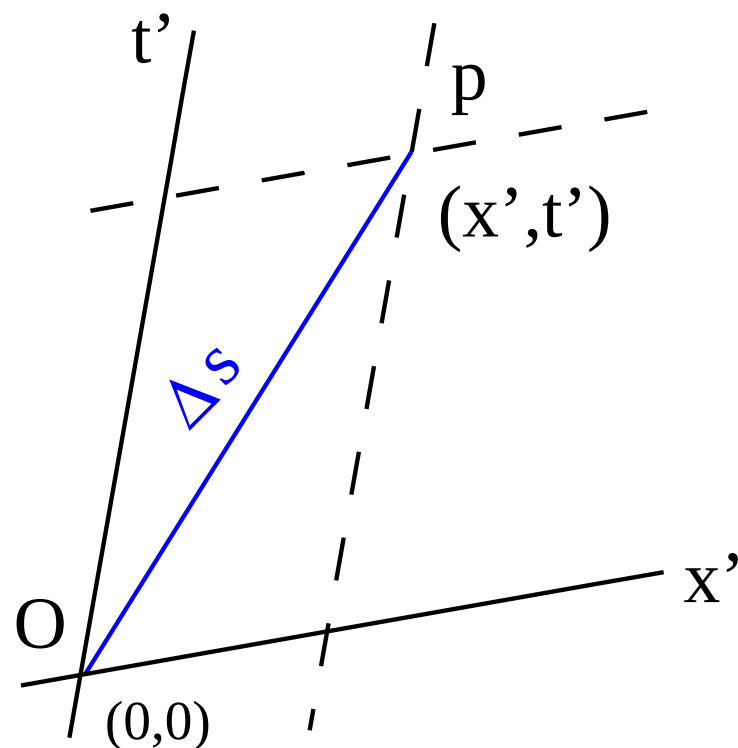
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$p$  at  $(x, t)$ , invariant interval from observer at  $O$  is  $\Delta s$   
 where  $(\Delta s)^2 = -(\Delta t)^2 + (\Delta x)^2$

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where velocity  $\beta := v/c \equiv v = \tanh \phi$



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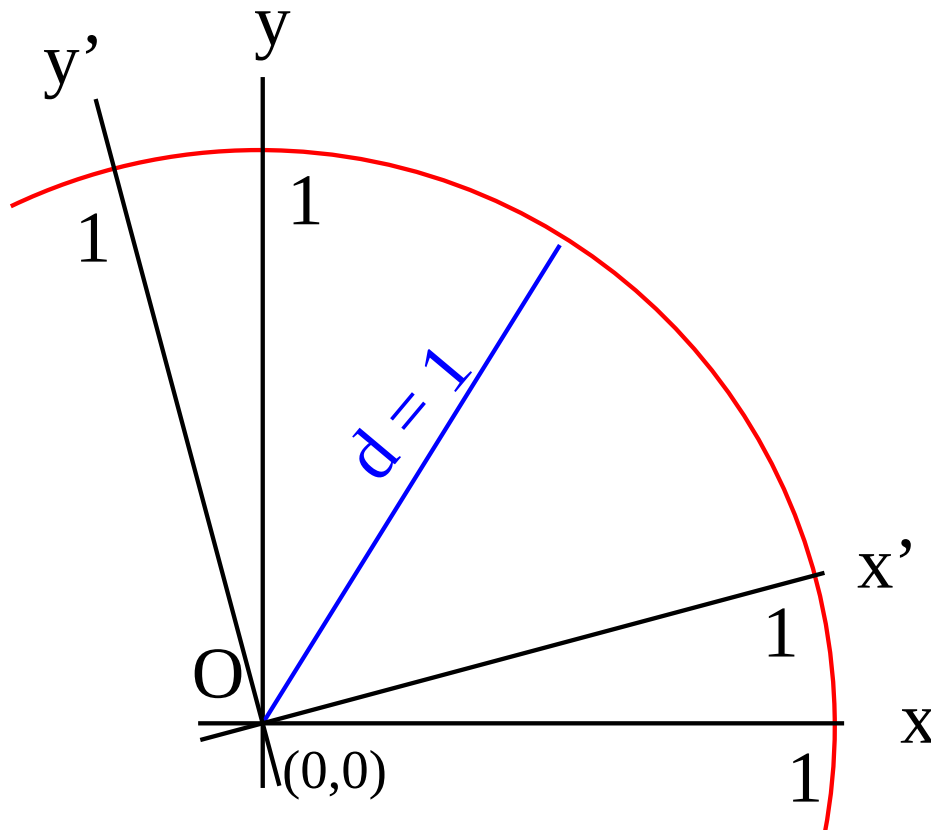
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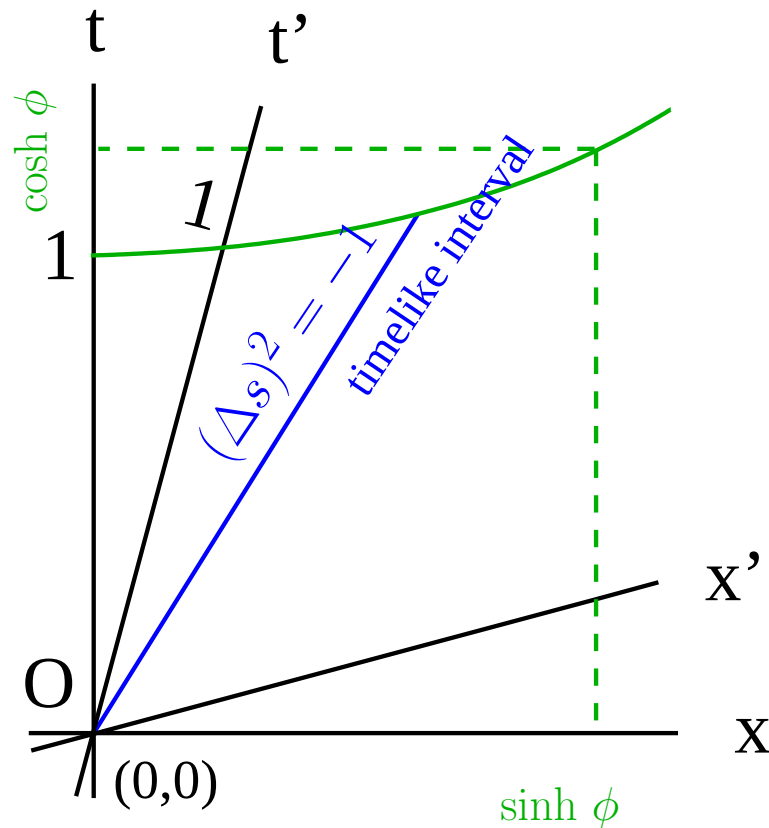


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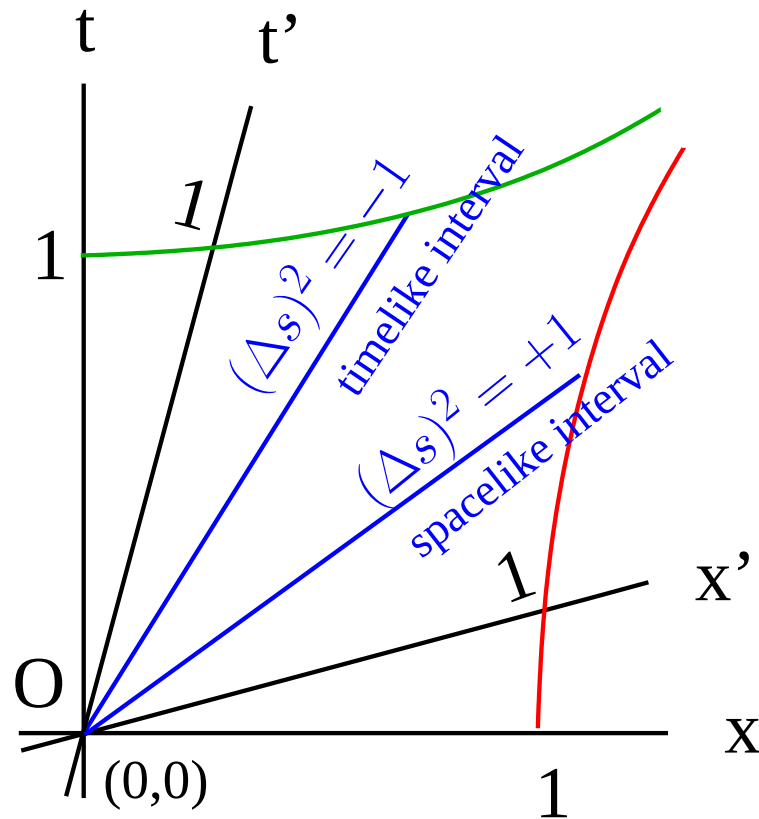
constant interval  
 $(\Delta s)^2 = -1$  from  $(0, 0)$

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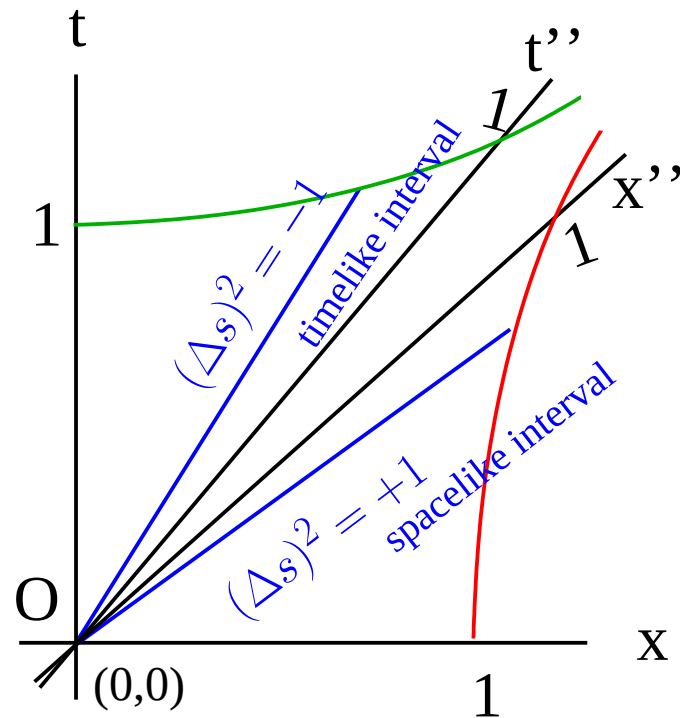
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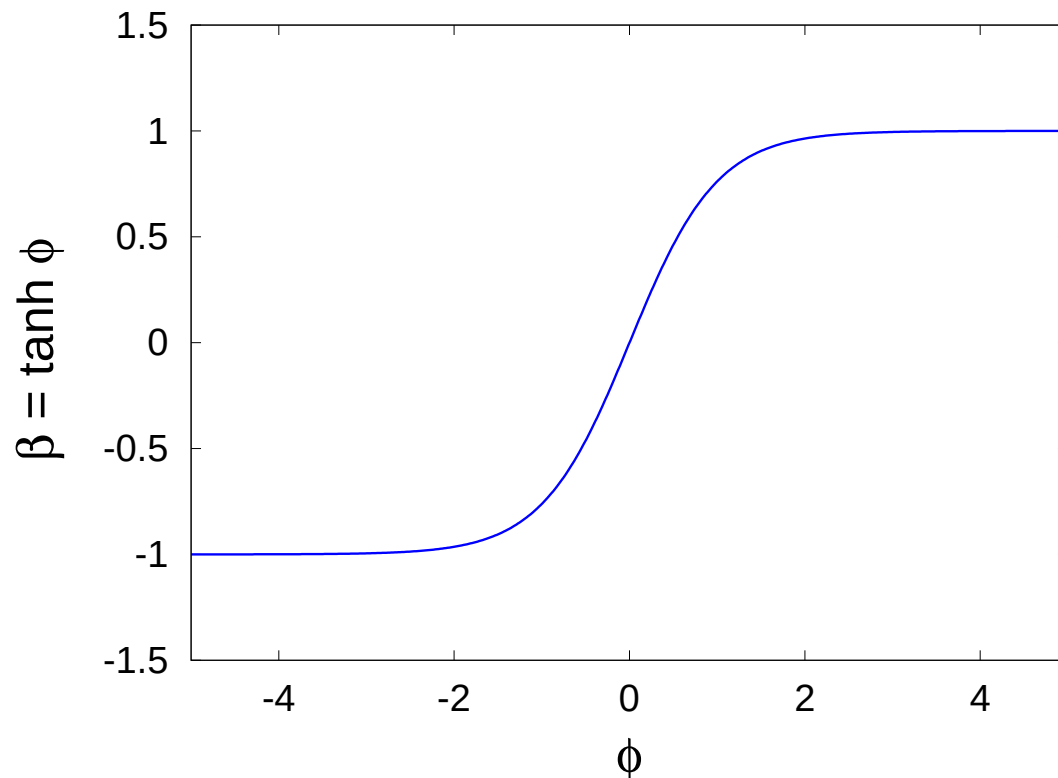


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- w:Michelson-Morley experiment (1887)

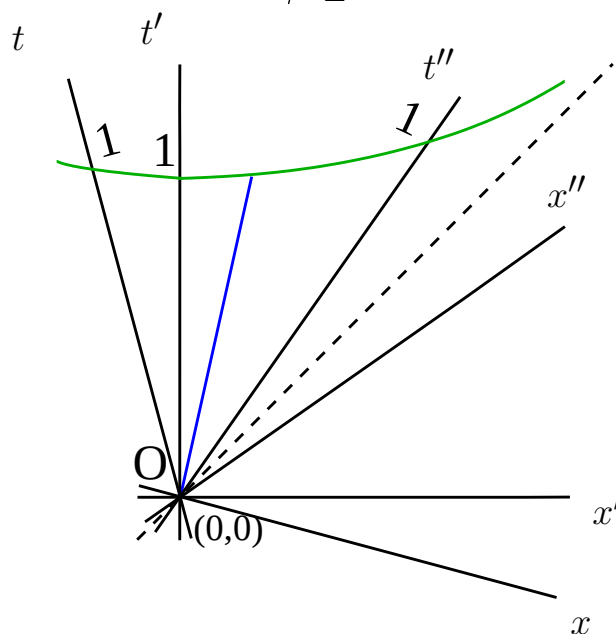


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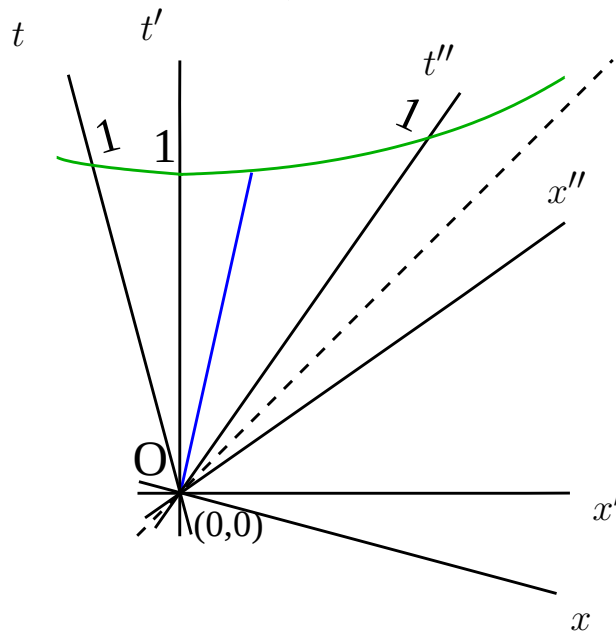


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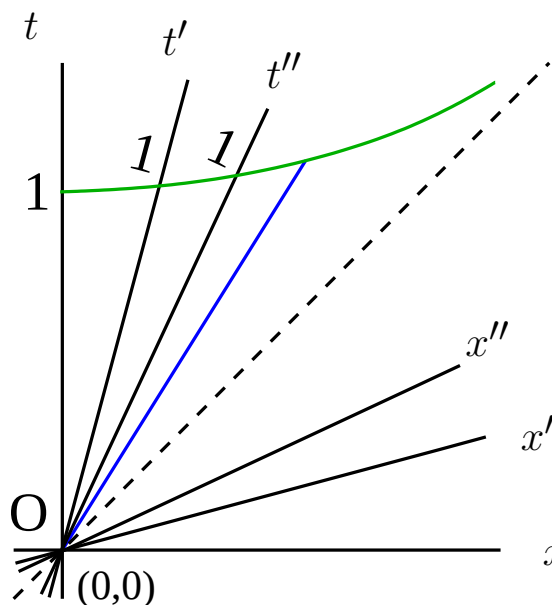


$$\begin{pmatrix} x'' \\ t'' \end{pmatrix} = \begin{pmatrix} \cosh \phi_2 & -\sinh \phi_2 \\ -\sinh \phi_2 & \cosh \phi_2 \end{pmatrix} \begin{pmatrix} x' \\ t' \end{pmatrix}$$

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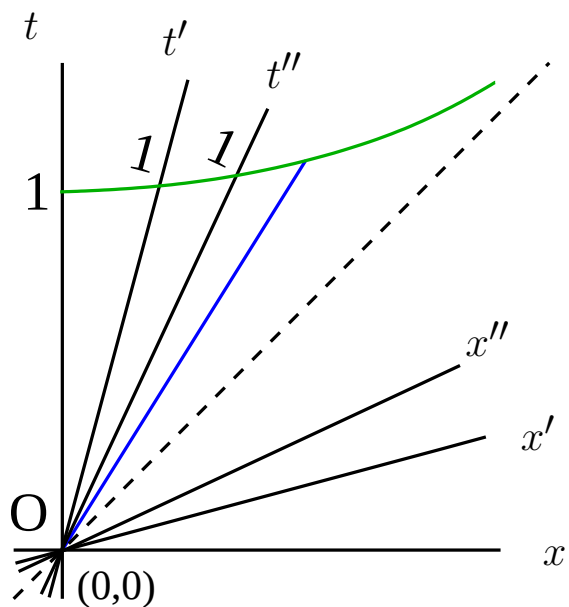
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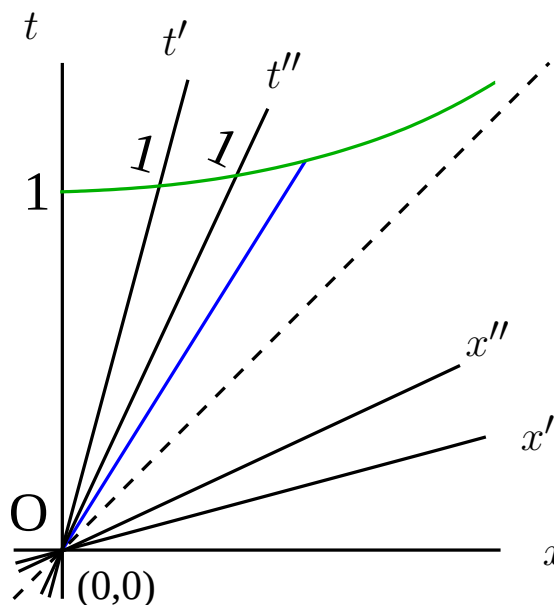
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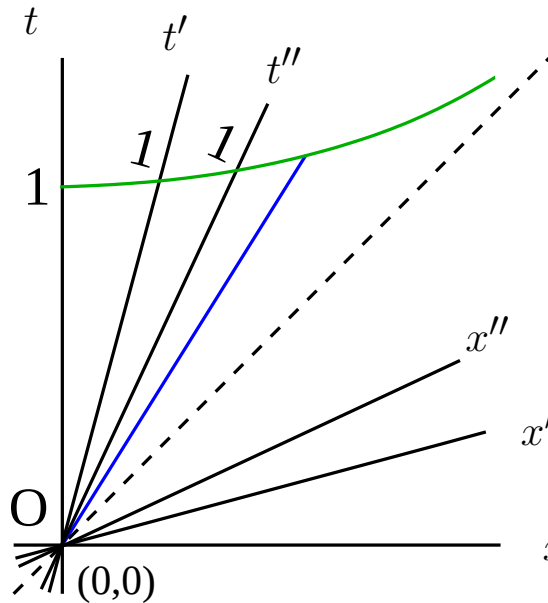


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but  $\Lambda(\phi_2)\Lambda(\phi_1) = \Lambda(\phi_1 + \phi_2)$

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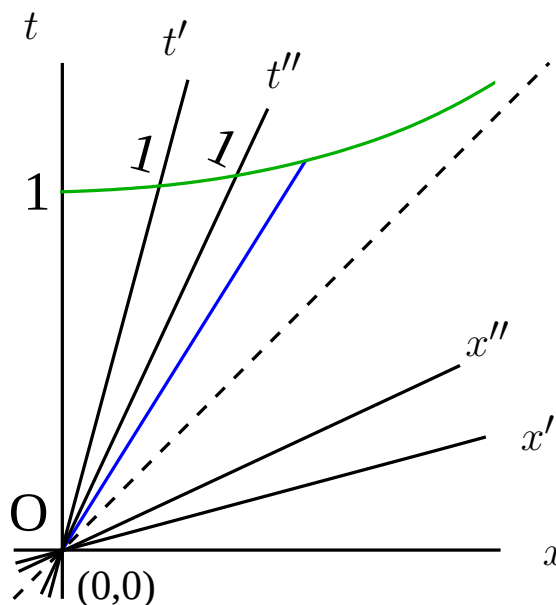
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but  $\Lambda(\phi_2)\Lambda(\phi_1) = \Lambda(\phi_1 + \phi_2)$

cf. rotation  $\theta_1$  “plus” rotation  $\theta_2 =$  rotation  $\theta_1 + \theta_2$

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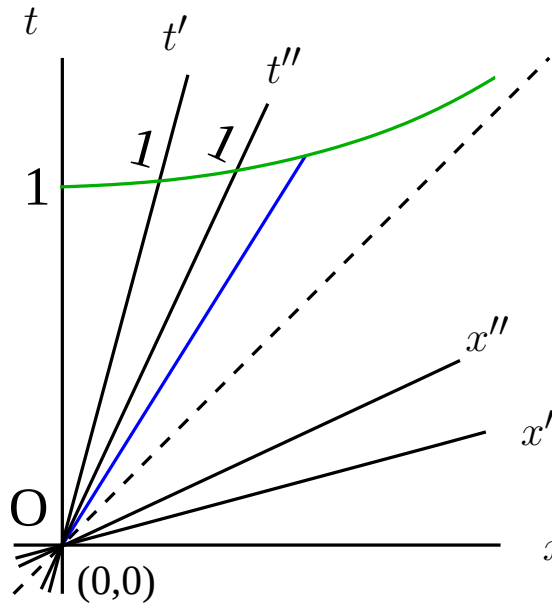
but  $\Lambda(\phi_2)\Lambda(\phi_1) = \Lambda(\phi_1 + \phi_2)$

so  $\beta_3 = \tanh(\phi_1 + \phi_2)$



# SR: adding velocities

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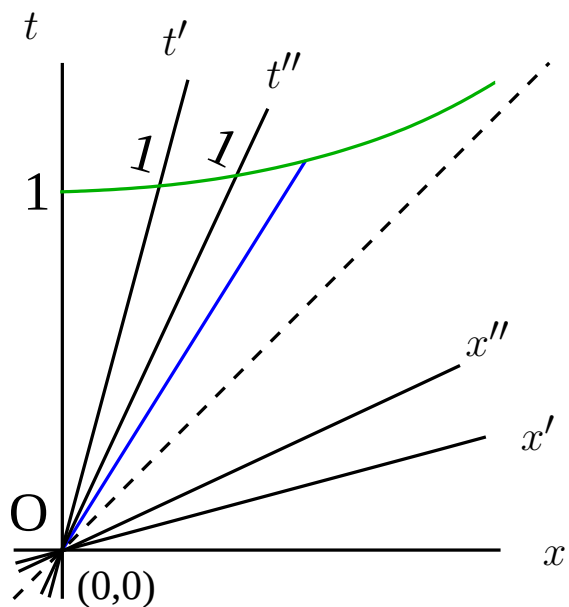
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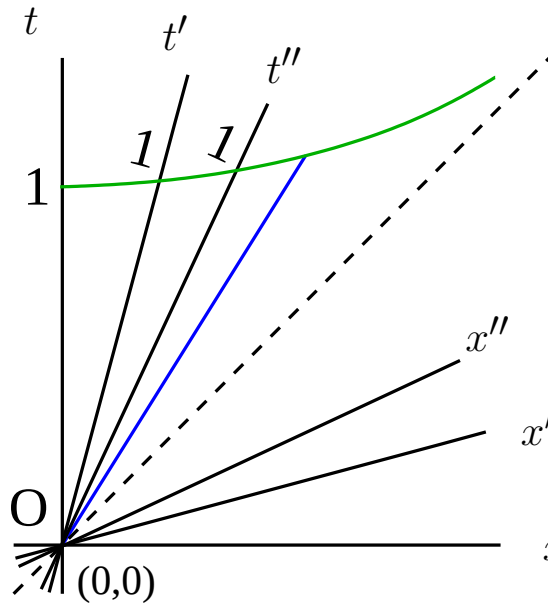
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but  $\Lambda(\phi_2)\Lambda(\phi_1) = \Lambda(\phi_1 + \phi_2)$

$$\text{so } \beta_3 = \tanh(\phi_1 + \phi_2) = \frac{\beta_1 + \beta_2}{1 + \beta_1\beta_2} = \frac{0.1 + 0.5}{1 + 0.1 \times 0.5} \approx 0.57$$

# SR: Lorentz factor



$\Lambda$ : alternative to hyperbolic trig functions



# SR: Lorentz factor

$\Lambda$ : alternative to hyperbolic trig functions

$$\Lambda(\phi) := \begin{pmatrix} \cosh \phi & -\sinh \phi \\ -\sinh \phi & \cosh \phi \end{pmatrix}$$
$$\cosh \phi := \frac{e^\phi + e^{-\phi}}{2}$$
$$\sinh \phi := \frac{e^\phi - e^{-\phi}}{2}$$

w:hyperbolic function

# SR: Lorentz factor



$\Lambda$ : alternative to hyperbolic trig functions

$$\Lambda(\beta) := \begin{pmatrix} \gamma & -\beta\gamma \\ -\beta\gamma & \gamma \end{pmatrix}$$

$$\beta = \tanh \phi$$

$$\gamma := (1 - \beta^2)^{-1/2} =$$

**Lorentz factor**

$$\gamma = \cosh \phi$$

$$\beta\gamma = \sinh \phi$$



# SR: model summary



Minkowski spacetime: draw a correct diagram



# SR: model summary



Minkowski spacetime: draw a correct diagram

Lorentz transformation (boost)  $\Lambda(\phi)$  or  $\Lambda(\beta)$





# SR: model summary



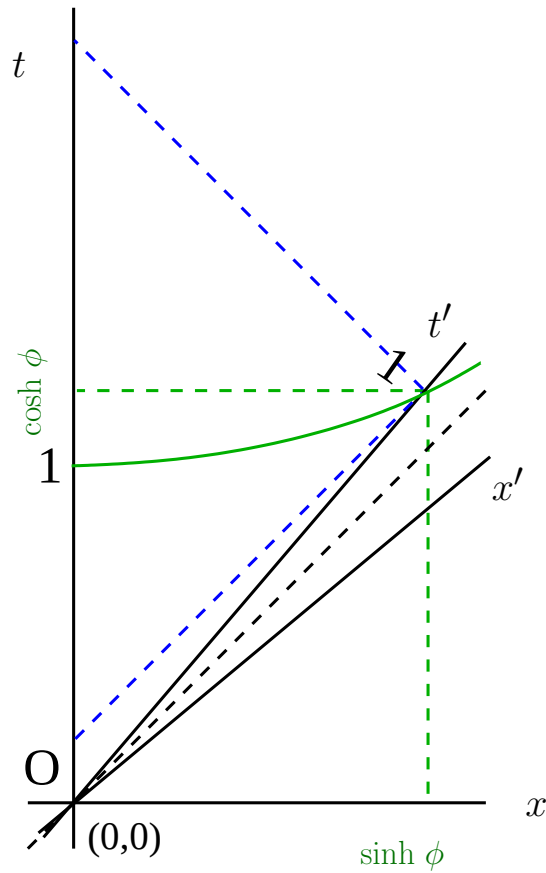
Minkowski spacetime: draw a correct diagram

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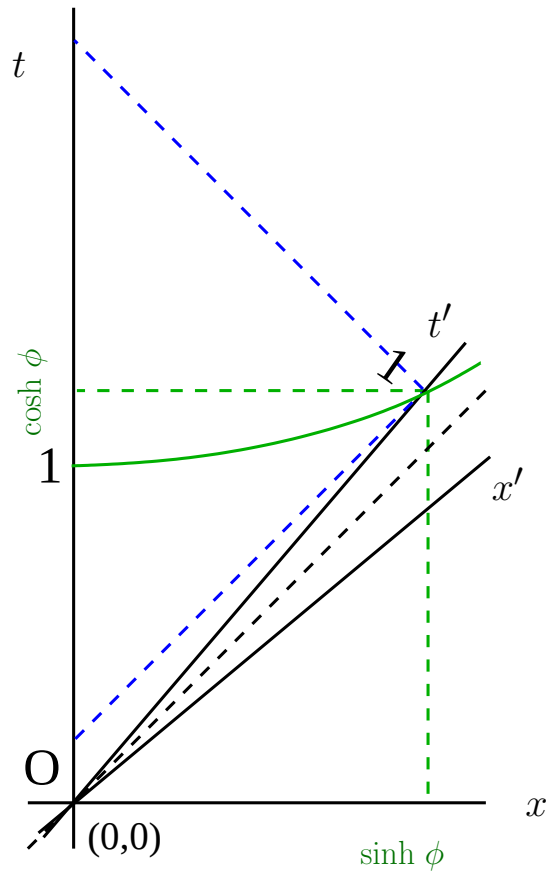
refuse the assumption of absolute simultaneity (time)



# SR: worldline time dilation

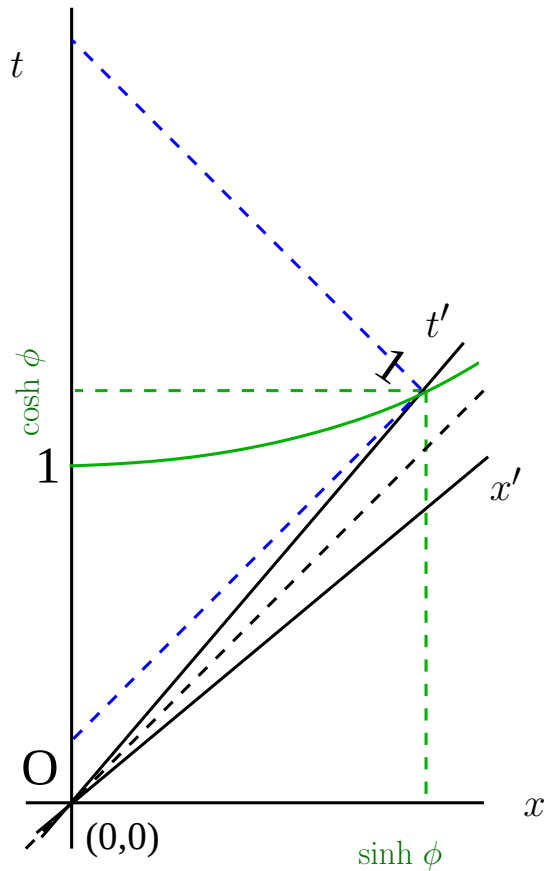


# SR: worldline time dilation



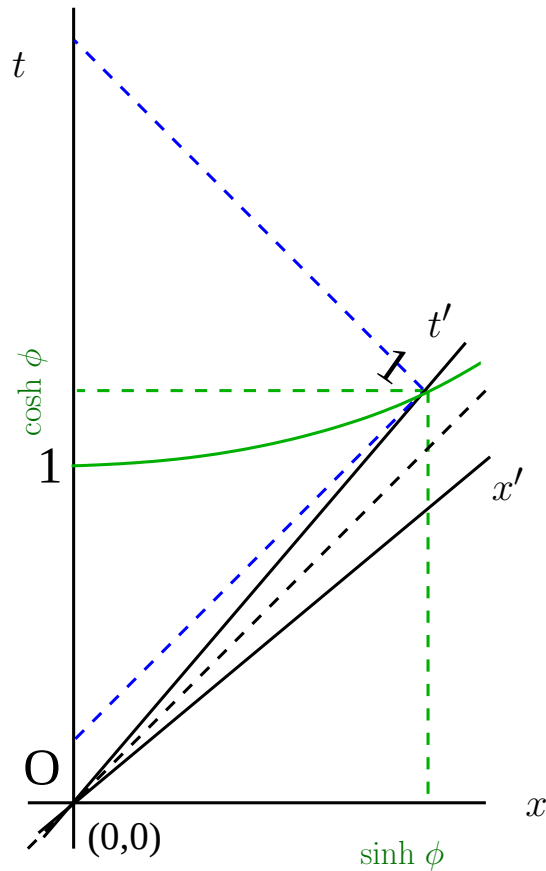
$$\cosh \phi \equiv \gamma$$

# SR: worldline time dilation



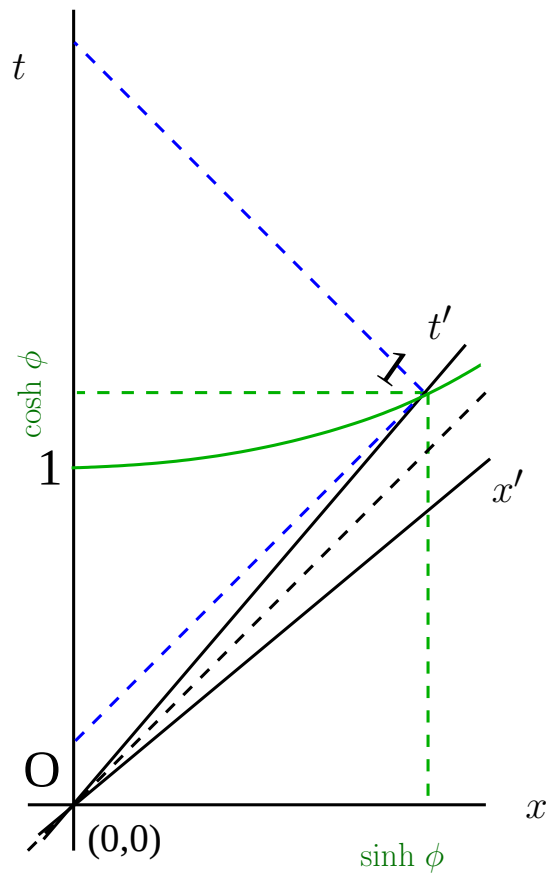
$$\cosh \phi \equiv \gamma \equiv \frac{1}{\sqrt{1-\beta^2}}$$

# SR: worldline time dilation



$$\cosh \phi \equiv \gamma \equiv \frac{1}{\sqrt{1-\beta^2}} > 1$$

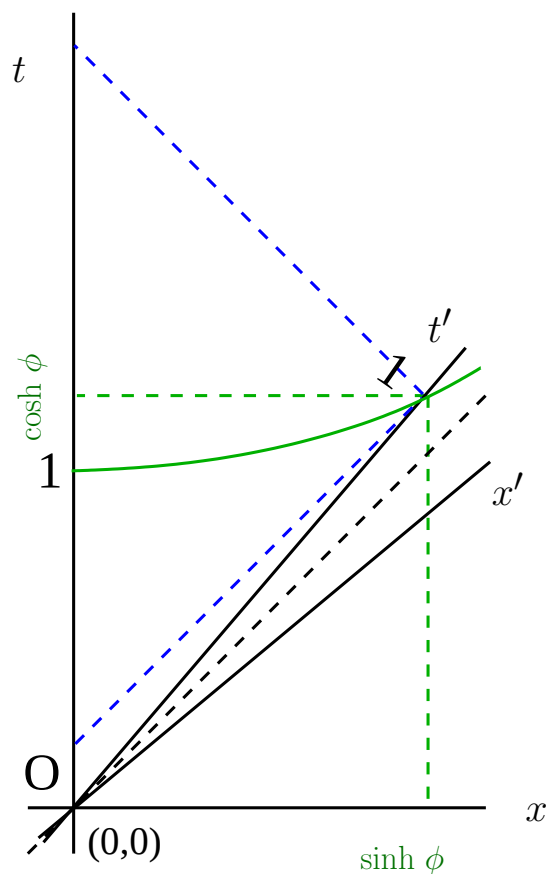
# SR: worldline time dilation



worldline “time dilation”

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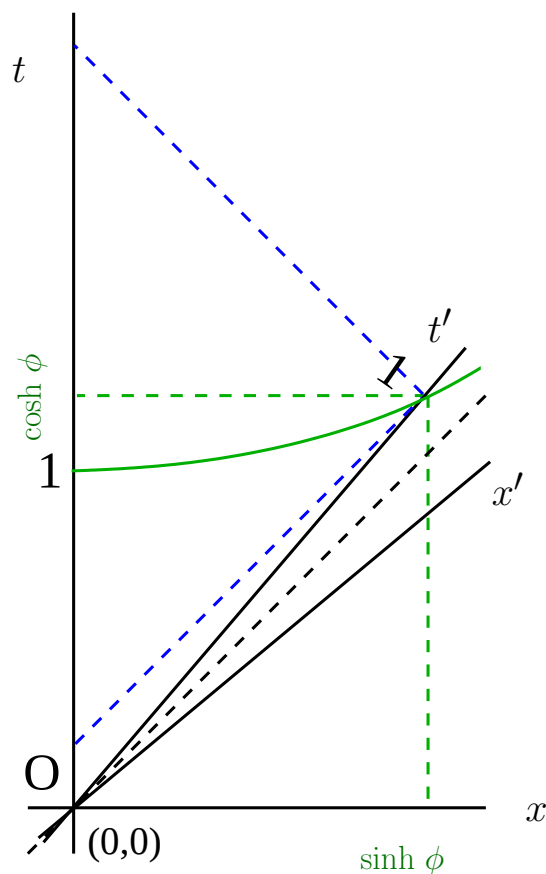


worldline “time dilation”

muons: mean lifetime  
2197 ns  $\ll$  15 km

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# SR: worldline time dilation



worldline “time dilation”

**muons:** mean lifetime  
2197 ns  $\ll$  15 km

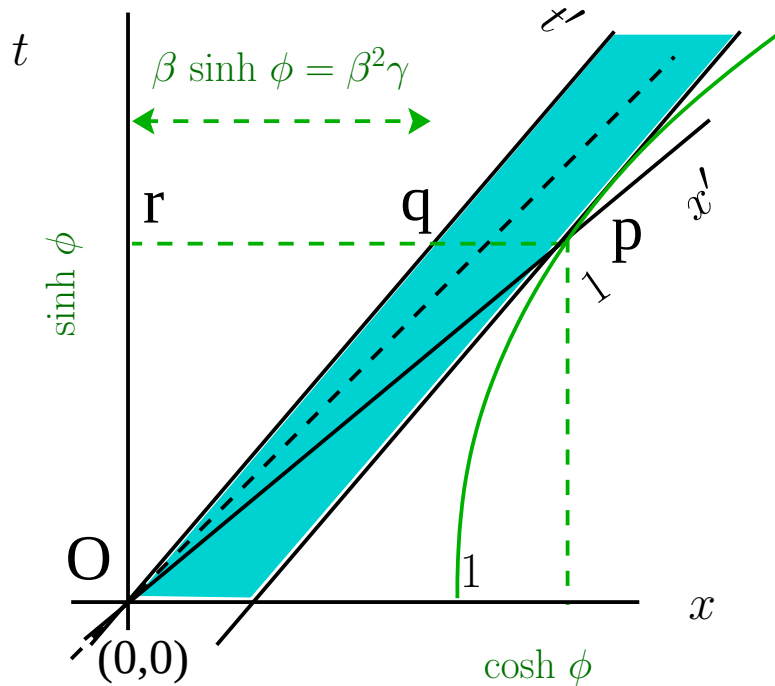
time dilation  $\Rightarrow$  muons  
can hit the ground

$$\cosh \phi \equiv \gamma \equiv \frac{1}{\sqrt{1-\beta^2}} > 1$$





# SR: worldsheet space contraction

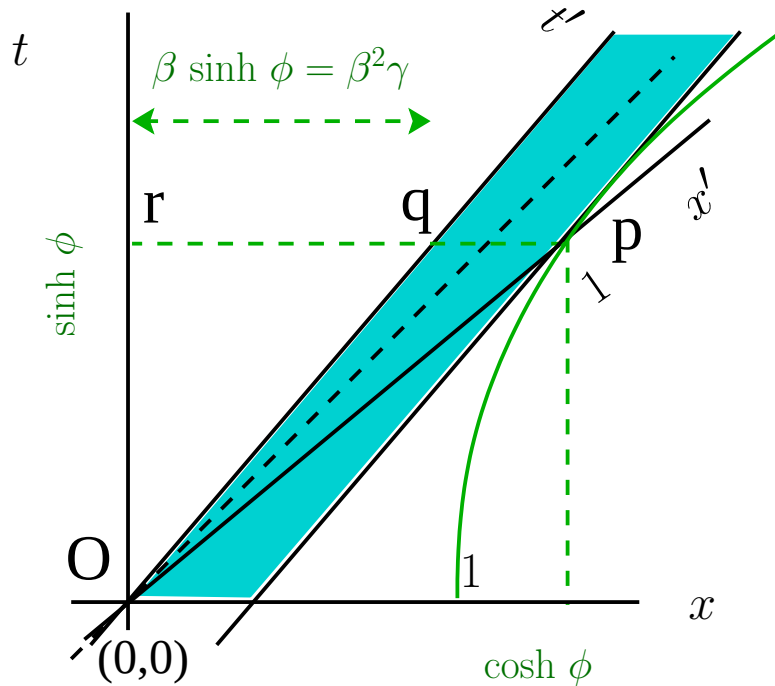


$$\sqrt{\Delta s^2(q, p)} =$$



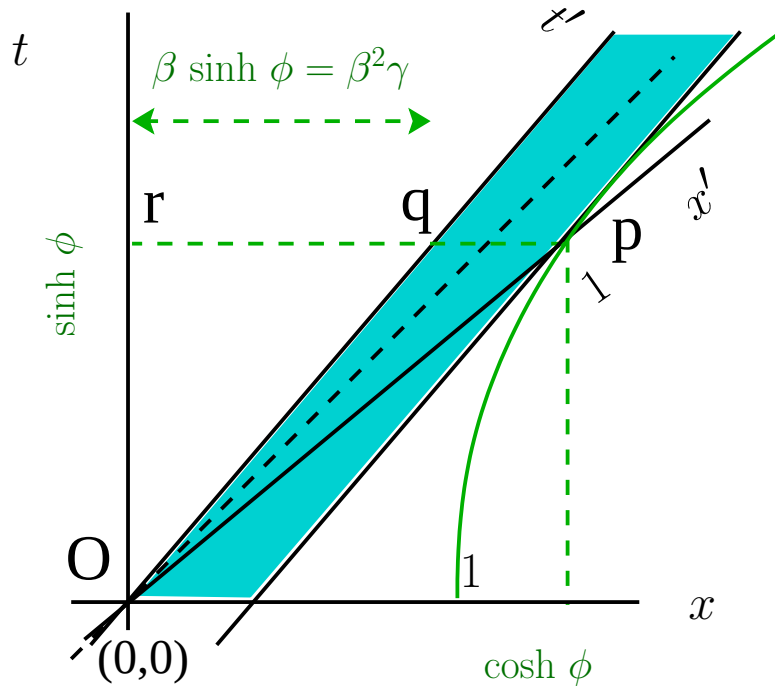
▽ – p.11

# SR: worldsheet space contraction



$$\sqrt{\Delta s^2(q, p)} = \gamma - \beta\beta\gamma$$

# SR: worldsheet space contraction

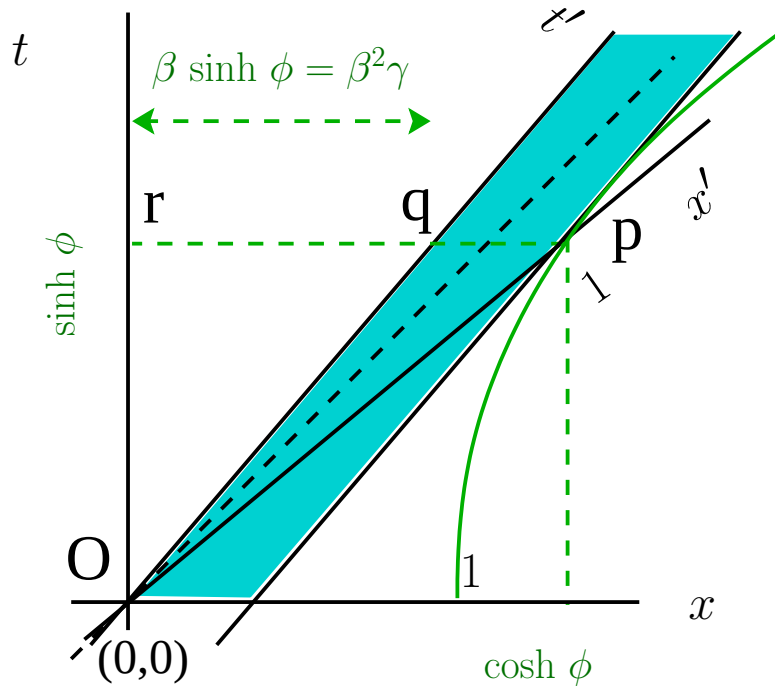


$$\sqrt{\Delta s^2(q, p)} = (1 - \beta^2)\gamma$$



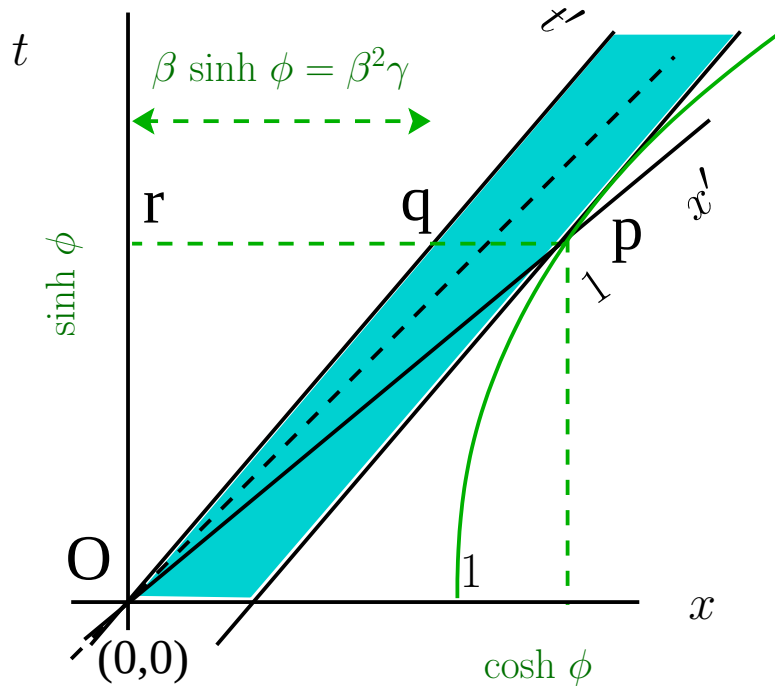
▽ – p.11

# SR: worldsheet space contraction



$$\sqrt{\Delta s^2(q, p)} = (1 - \beta^2)^{1/2}$$

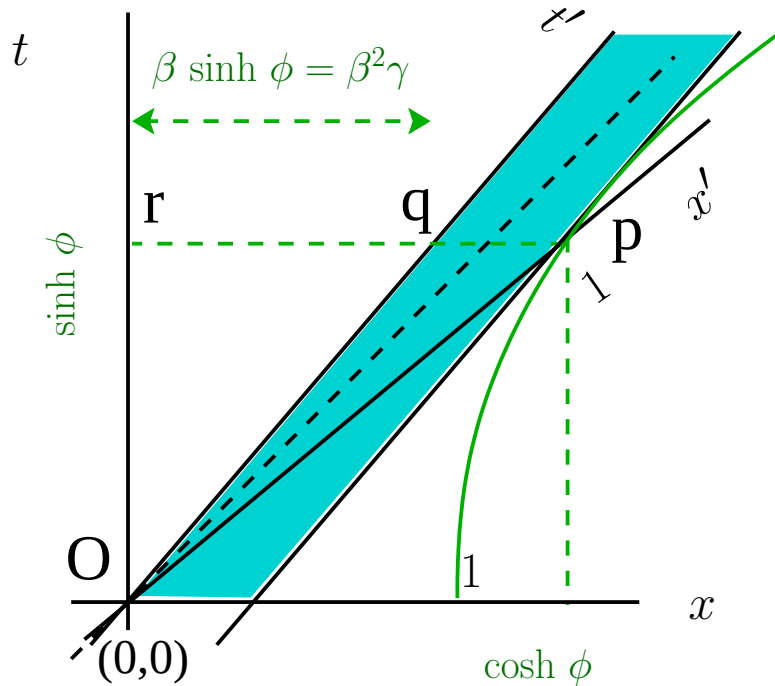
# SR: worldsheet space contraction



$$\sqrt{\Delta s^2(q, p)} = \gamma^{-1} < 1$$

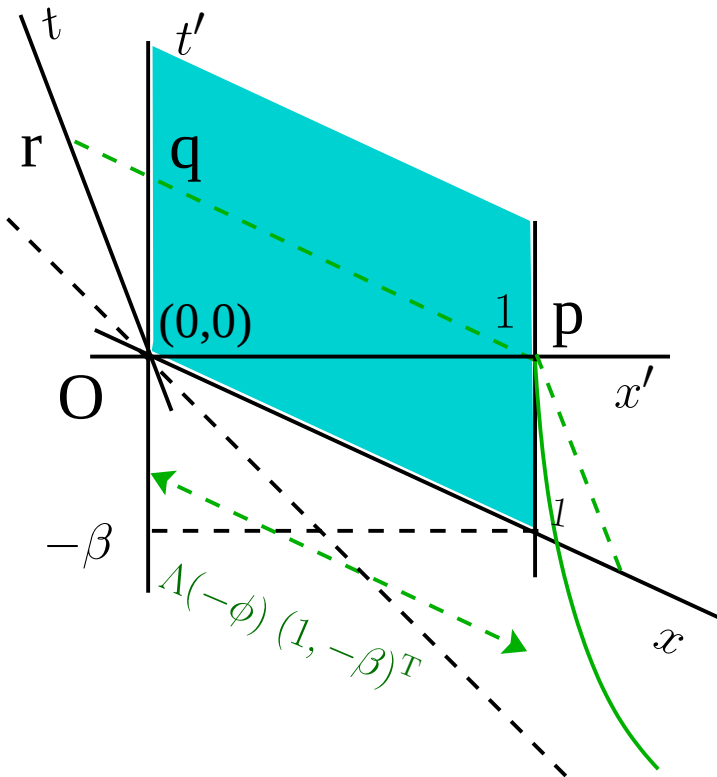


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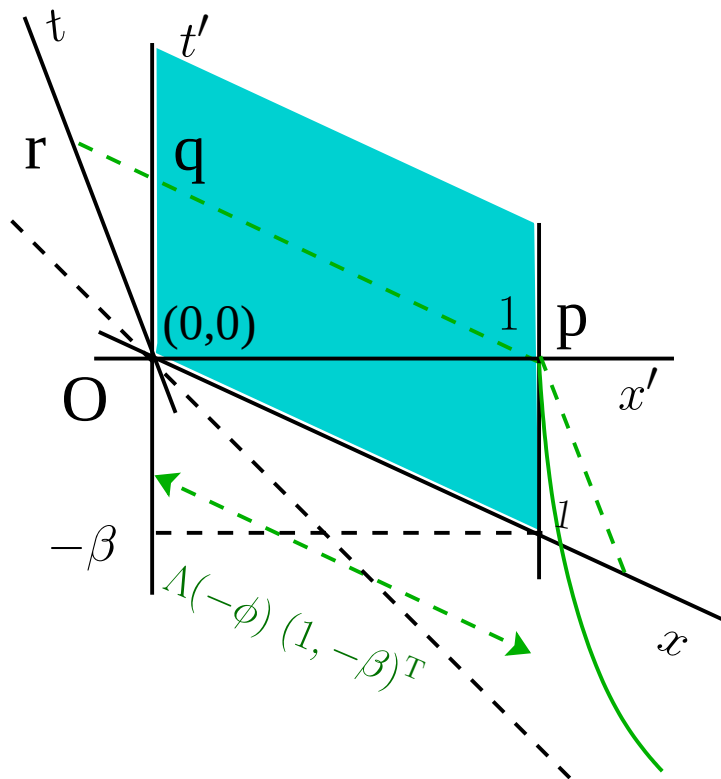


$$\sqrt{\Delta s^2(q, p)} = \gamma^{-1} < 1 \quad \text{worldsheet "space contraction"}$$

# SR: worldsheet space contraction

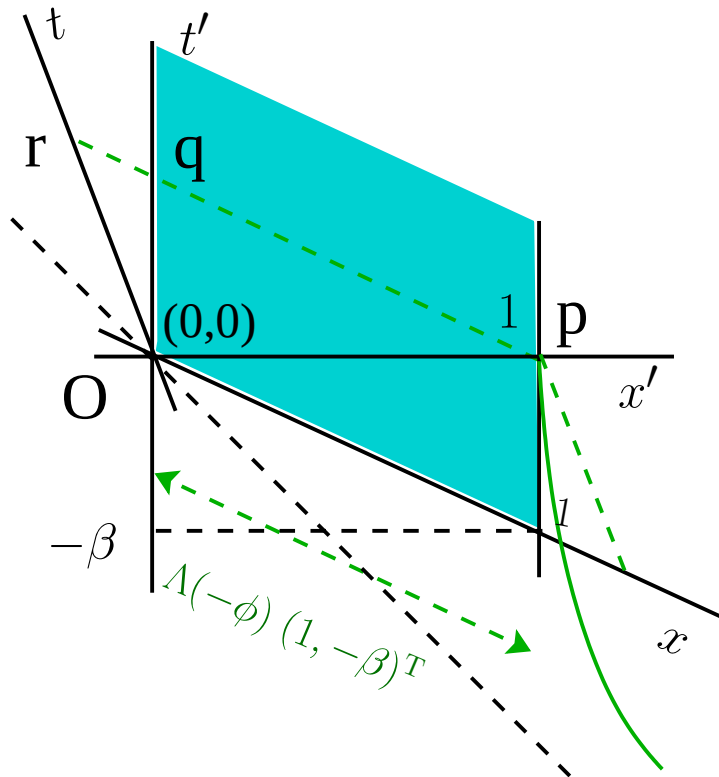


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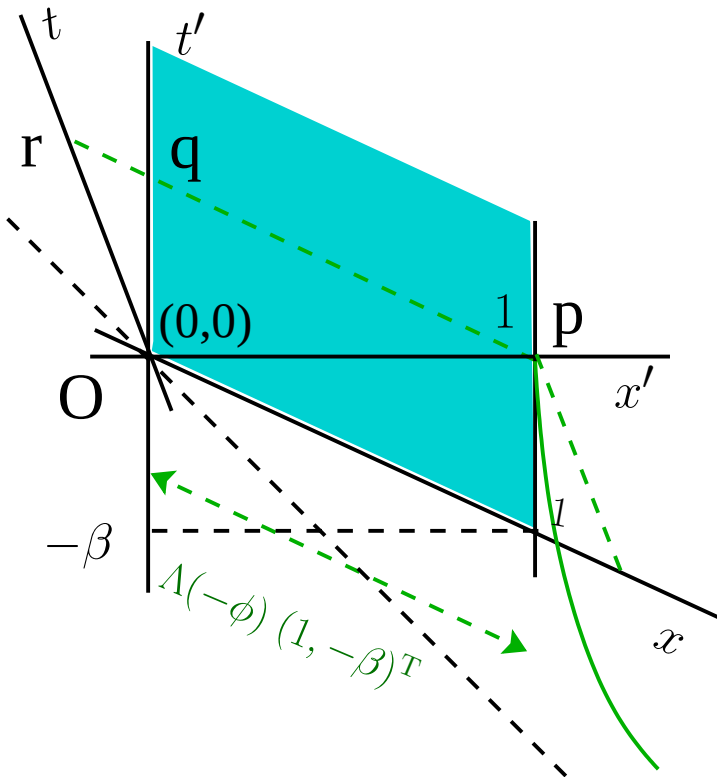
$$\Lambda^{-1} \begin{pmatrix} 1 \\ -\beta \end{pmatrix} = \begin{pmatrix} \cosh \phi - \beta \sinh \phi \\ \sinh \phi - \beta \cosh \phi \end{pmatrix}$$

# SR: worldsheet space contraction



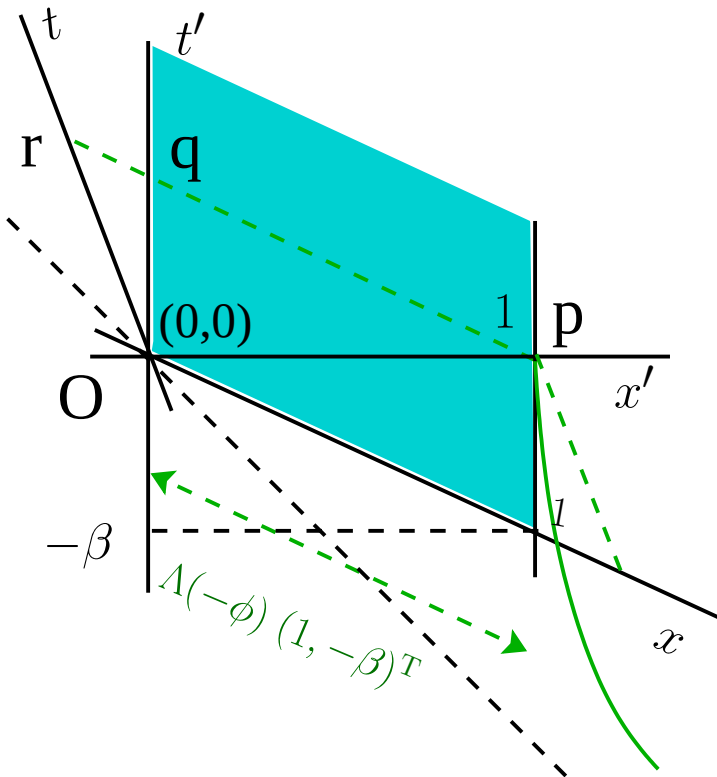
$$\Lambda^{-1} \begin{pmatrix} 1 \\ -\beta \end{pmatrix} = \begin{pmatrix} \cosh \phi (1 - \beta^2) \\ \cosh \phi (\tanh \phi - \tanh \phi) \end{pmatrix}$$

# SR: worldsheet space contraction



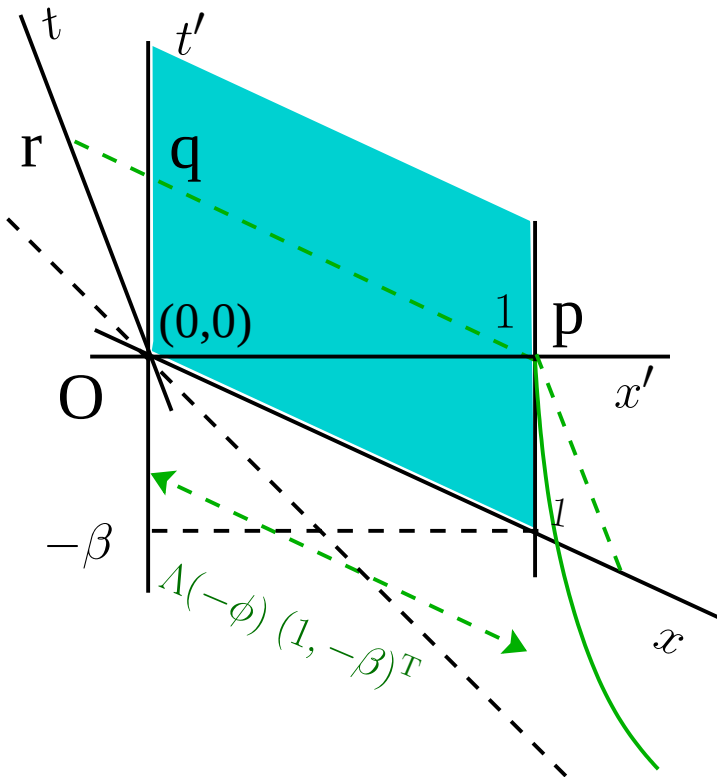
$$\Lambda^{-1} \begin{pmatrix} 1 \\ -\beta \end{pmatrix} = \begin{pmatrix} \gamma(1 - \beta^2) \\ 0 \end{pmatrix}$$

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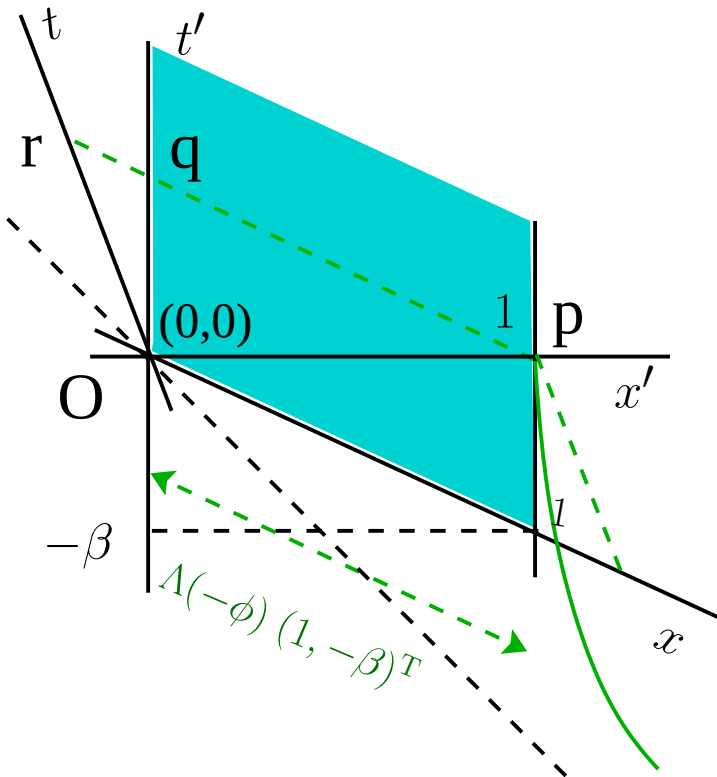
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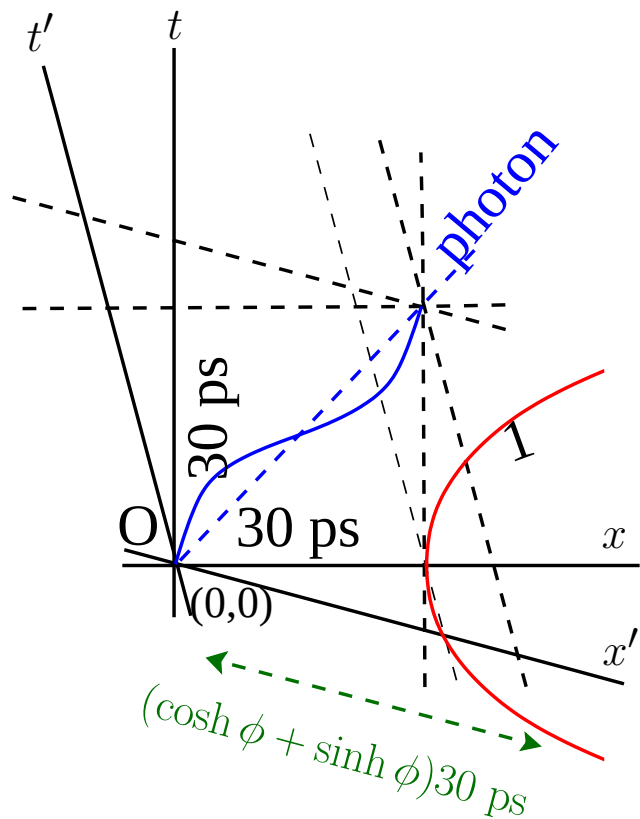
# SR: worldsheet space contraction



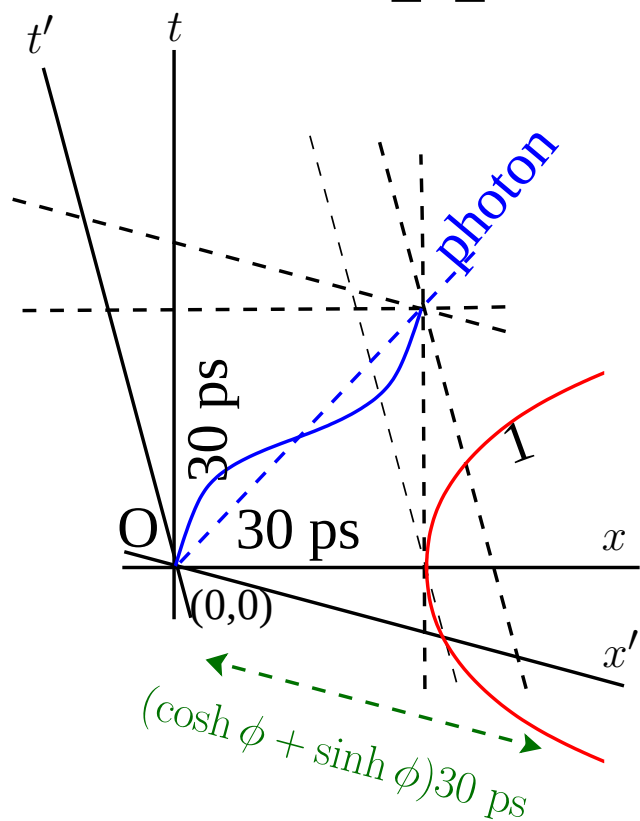
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# SR: Doppler shift

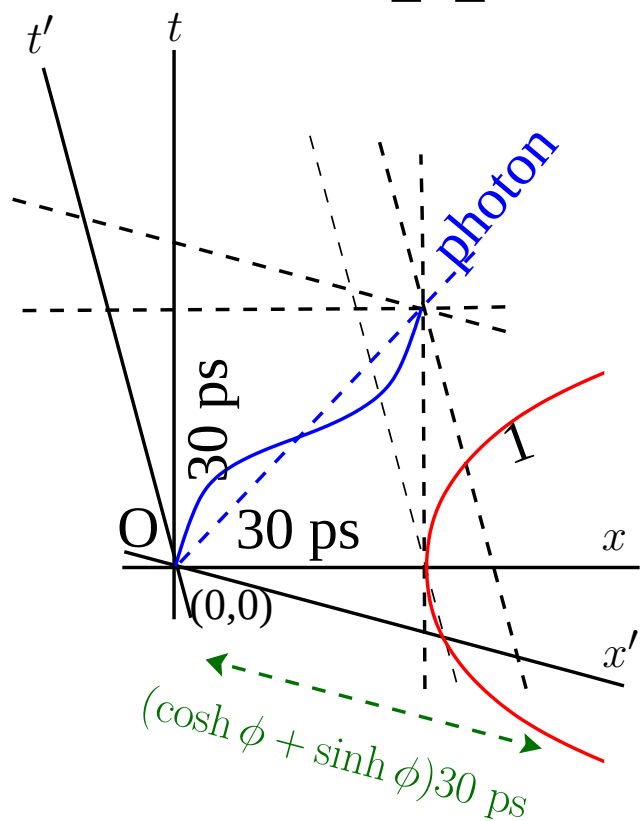


# SR: Doppler shift



see photon worldline calculation

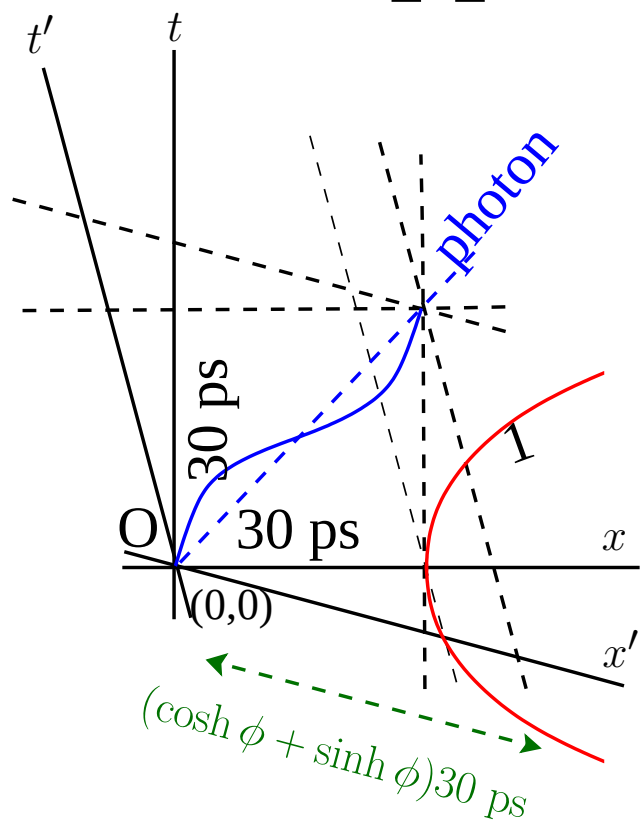
# SR: Doppler shift



see photon worldline calculation

$$x' = (\cosh \phi + \sinh \phi)t$$

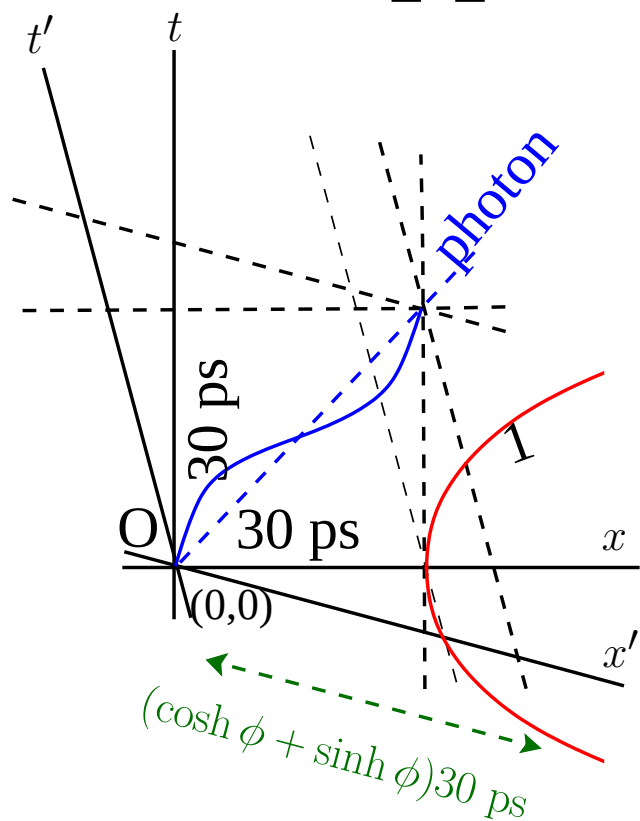
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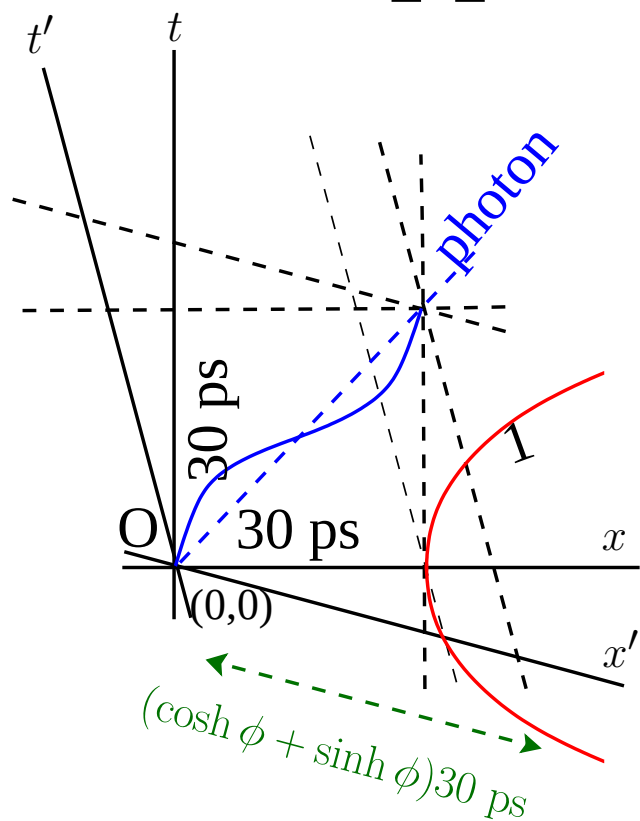
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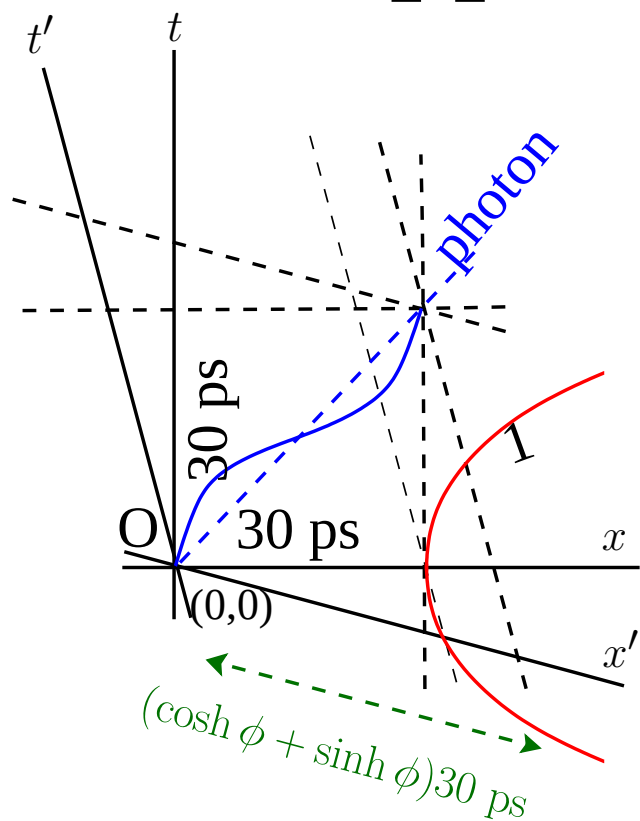
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see photon worldline calculation

$$x'/x = \gamma + \beta\gamma$$

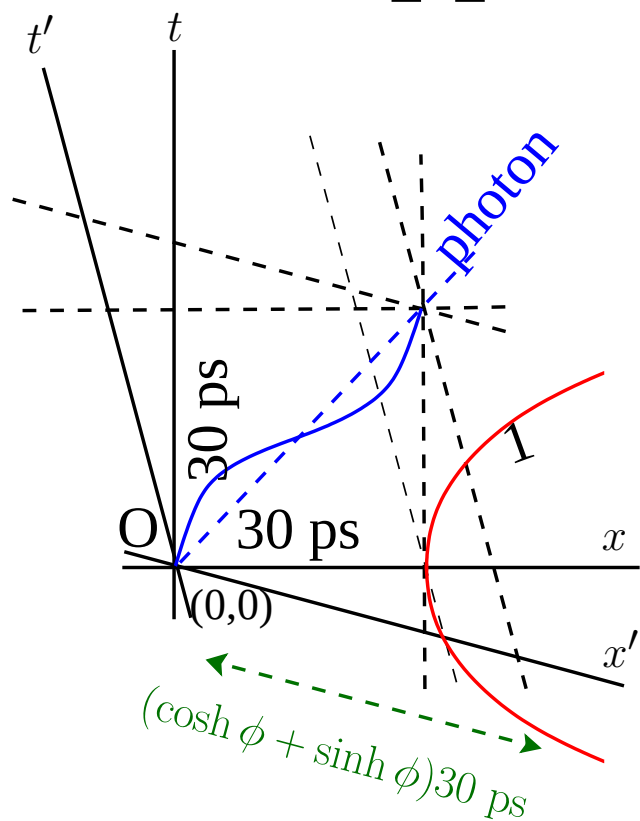
# SR: Doppler shift



see photon worldline calculation

$$x'/x = \gamma(1 + \beta)$$

# SR: Doppler shift

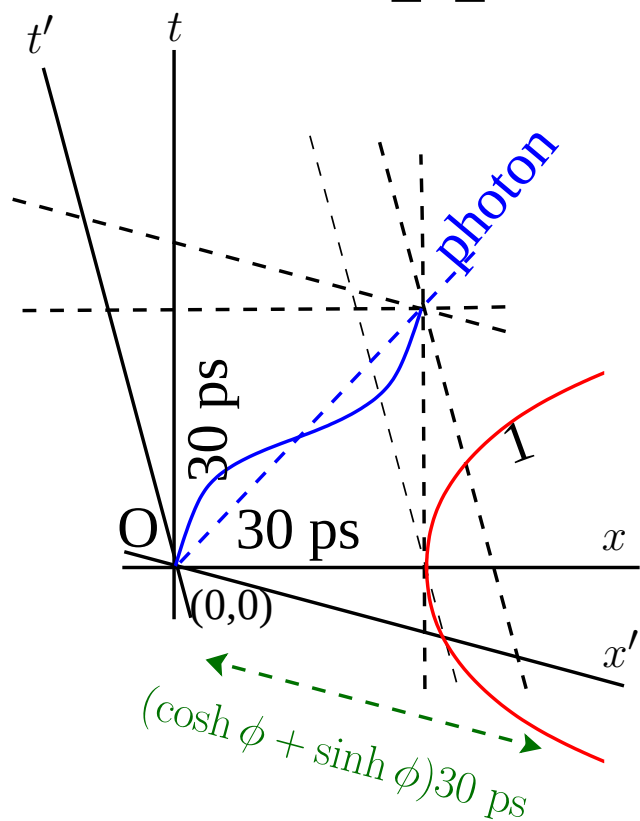


see photon worldline calculation

$$x'/x = \gamma(1 + \beta) = \frac{1+\beta}{\sqrt{1-\beta^2}}$$



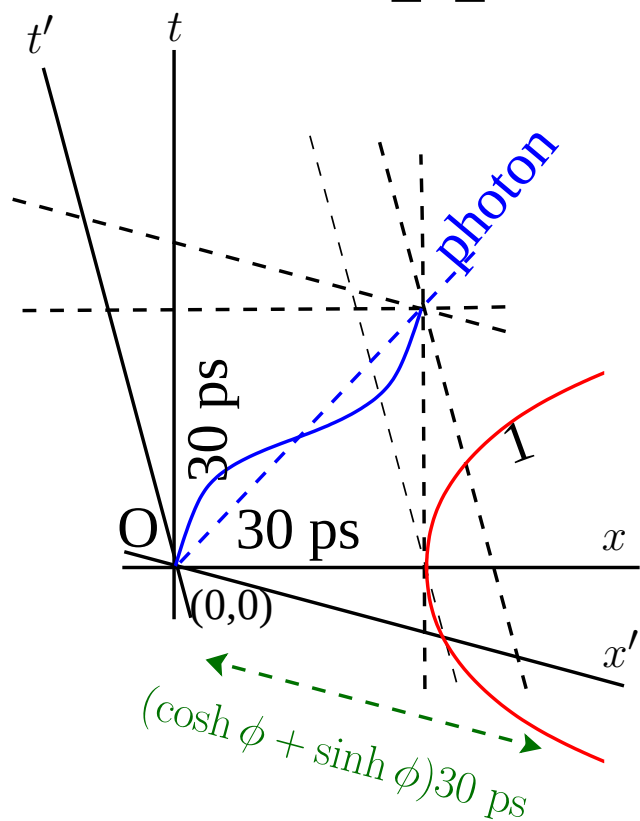
# SR: Doppler shift



see photon worldline calculation

$$x'/x = \gamma(1 + \beta) = \frac{\sqrt{(1+\beta)^2}}{\sqrt{(1-\beta)(1+\beta)}}$$

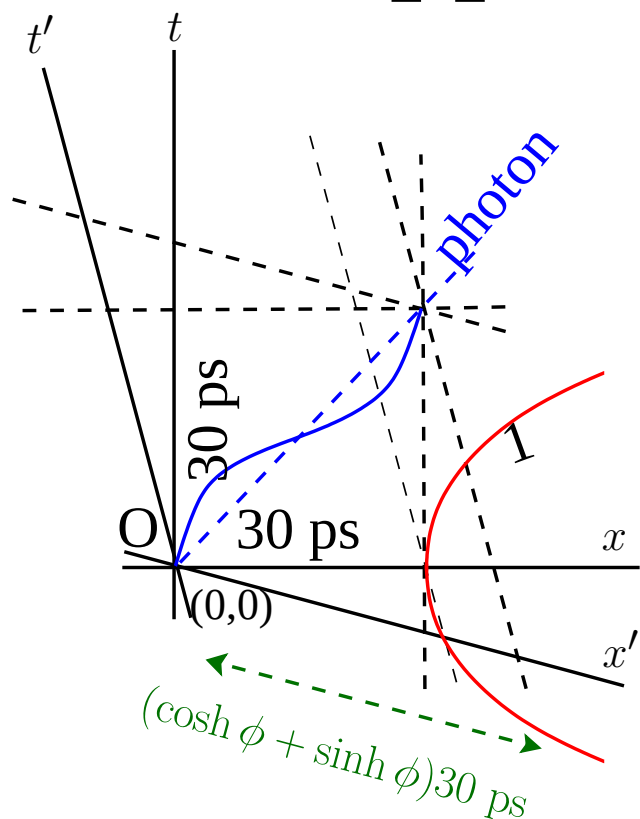
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see photon worldline calculation

$$x'/x = \gamma(1 + \beta) = \sqrt{\frac{1+\beta}{1-\beta}}$$

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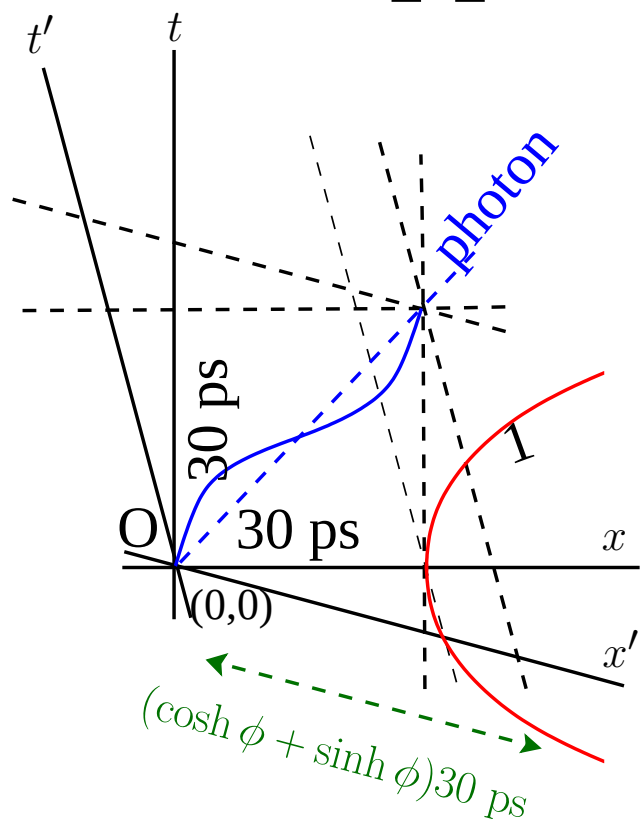


see photon worldline calculation

$$1 + z := \lambda' / \lambda = \gamma(1 + \beta) \equiv \sqrt{\frac{1+\beta}{1-\beta}}$$

redshift

# SR: Doppler shift



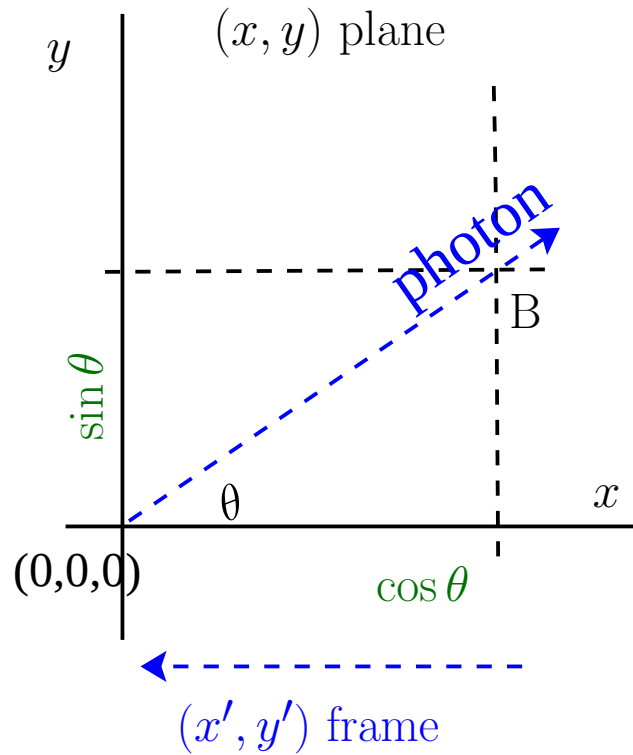
see photon worldline calculation

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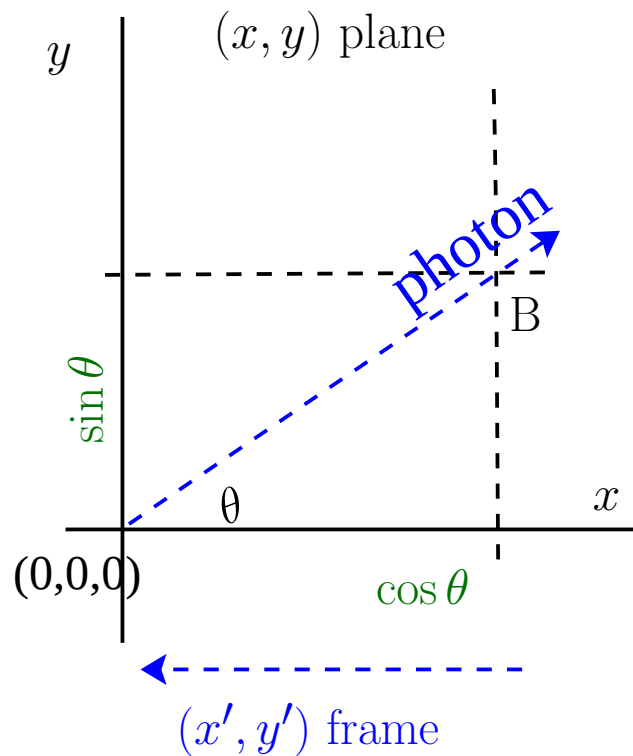
redshift

$\Rightarrow$  when  $\beta \ll 1$ ,  $z \approx \beta$

# SR: relativistic aberration



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event B:  $(x, y, t) = (\cos \theta, \sin \theta, 1)$

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$$\Lambda^{-1} \begin{pmatrix} \cos \theta \\ \sin \theta \\ 1 \end{pmatrix}$$



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event B:  $(x, y, t) = (\cos \theta, \sin \theta, 1)$

$$\Lambda^{-1} \begin{pmatrix} \cos \theta \\ \sin \theta \\ 1 \end{pmatrix} = \begin{pmatrix} \gamma & 0 & \beta\gamma \\ 0 & 1 & 0 \\ \beta\gamma & 0 & \gamma \end{pmatrix} \begin{pmatrix} \cos \theta \\ \sin \theta \\ 1 \end{pmatrix}$$

# SR: relativistic aberration

event B:  $(x, y, t) = (\cos \theta, \sin \theta, 1)$

$$\Lambda^{-1} \begin{pmatrix} \cos \theta \\ \sin \theta \\ 1 \end{pmatrix} = \begin{pmatrix} \gamma \cos \theta + \beta \gamma \\ \sin \theta \\ \beta \gamma \cos \theta + \gamma \end{pmatrix}$$

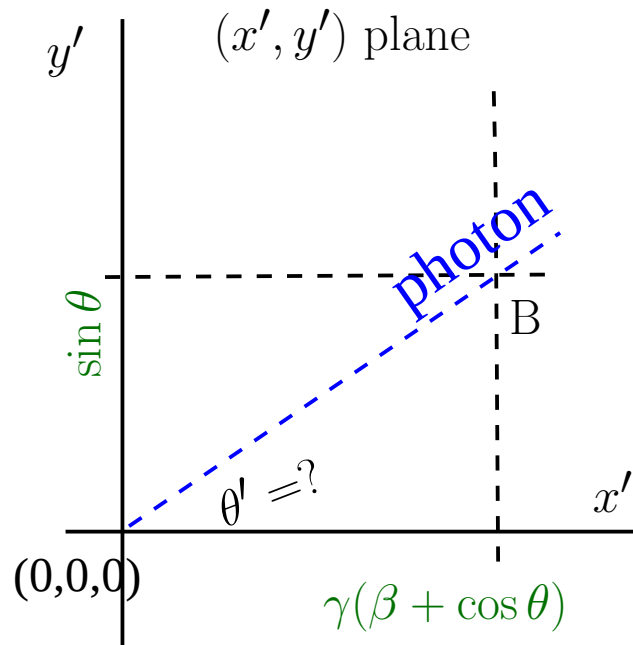
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event B:  $(x, y, t) = (\cos \theta, \sin \theta, 1)$

$$\Lambda^{-1} \begin{pmatrix} \cos \theta \\ \sin \theta \\ 1 \end{pmatrix} = \begin{pmatrix} \gamma(\cos \theta + \beta) \\ \sin \theta \\ \gamma(1 + \beta \cos \theta) \end{pmatrix}$$

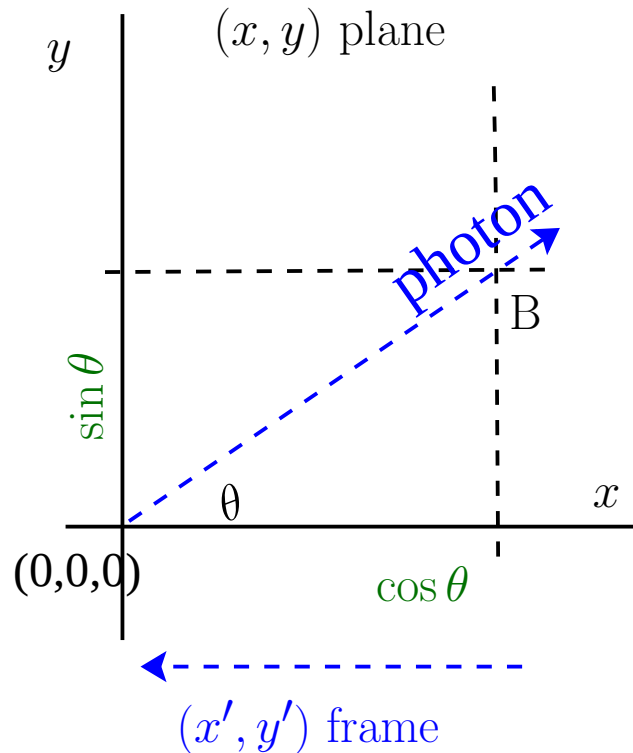
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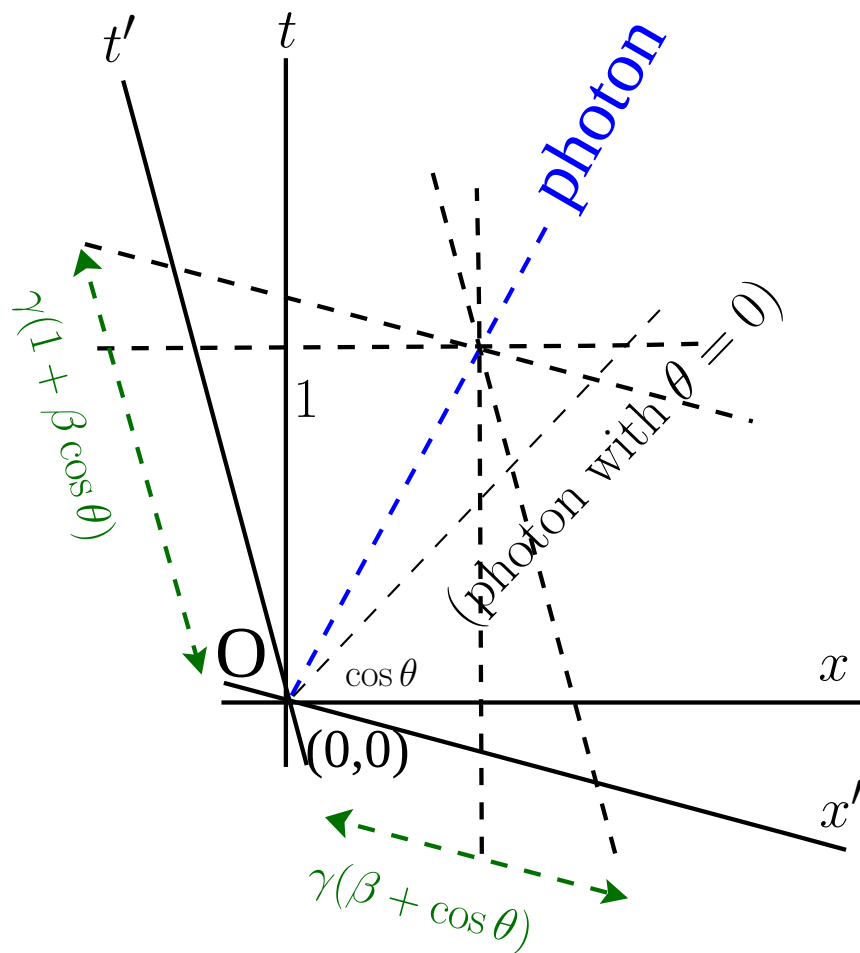
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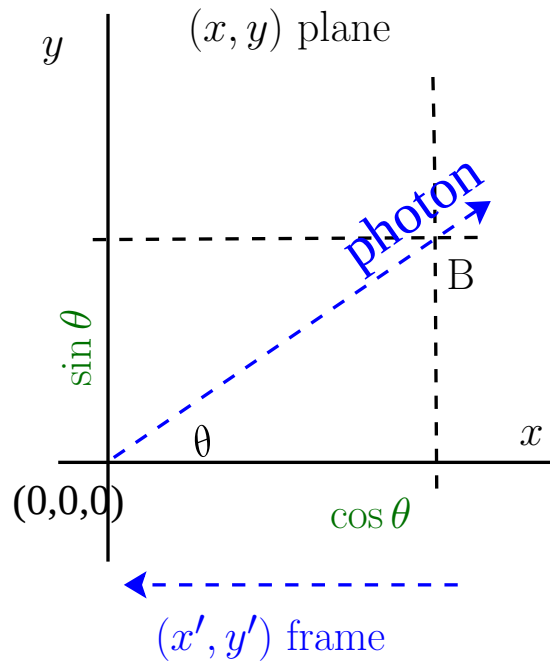
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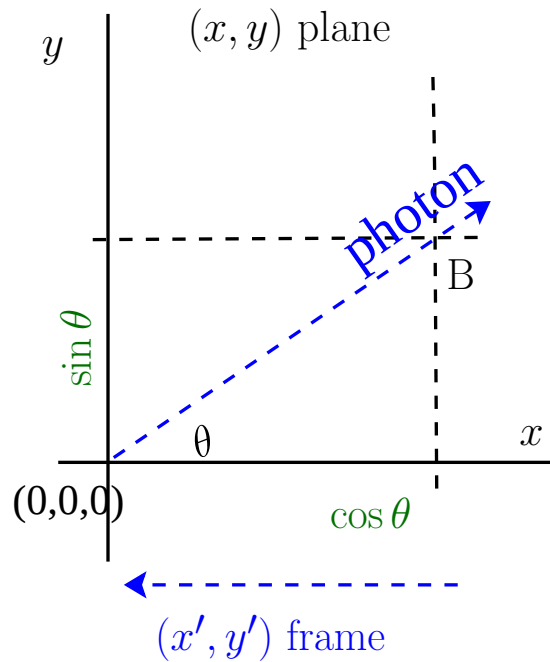
event B:  $(x, y, t) = (\cos \theta, \sin \theta, 1)$



$$\tan \theta' = \frac{\sin \theta}{\gamma(\beta + \cos \theta)}$$

# SR: relativistic aberration

event B:  $(x, y, t) = (\cos \theta, \sin \theta, 1)$



$$\tan \theta' = \frac{\sin \theta}{\gamma(\beta + \cos \theta)}$$

⇒ relativistic beaming, e.g. AGN jets



# SR: four-velocity, ...



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# GR:

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+ w:Riemannian geometry (and pseudo-Riemannian geometry)

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# GR:

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+ w:Intermediate treatment of tensors

# GR:

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# GR:

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# GR:

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# GR:

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+ w:Intermediate treatment of tensors

# GR:

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+ w:Intermediate treatment of tensors



# GR: maxima

+ maxima - component tensor packet ctensor

# GR: maxima

+ maxima - component tensor packet ctensor

# GR: maxima

+ maxima - component tensor packet ctensor

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+ maxima - component tensor packet ctensor

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# GR:

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+ w:Einstein field equations

# GR:

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# GR:

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+ w:Equivalence principle

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# GR:

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+ w:Schwarzschild metric

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# GR:

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w:Friedmann-Lemaître-Robertson-Walker metric



# GR:

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# GR:

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w:Friedmann-Lemaître-Robertson-Walker metric





# GR: an approximation method: ADM

+ w:ADM formalism



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# GR: a numerical method: cactus

+ Cactus - <http://cactuscode.org>



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